

Geometric Mechanics

Homework 3

Due on October 12

1. Consider the usual local coordinates (θ, φ) on the sphere $S^2 \subset \mathbb{R}^3$, defined by the parameterization $\phi :]0, \pi[\times]0, 2\pi[\rightarrow \mathbb{R}^3$ given by

$$\phi(\theta, \varphi) = (\sin \theta \cos \varphi, \sin \theta \sin \varphi, \cos \theta).$$

Recall that the Riemannian metric g induced in S^2 by the usual Euclidean metric in \mathbb{R}^3 is given in these coordinates by

$$g = d\theta \otimes d\theta + \sin^2 \theta d\varphi \otimes d\varphi.$$

- (a) Determine the Christoffel symbols for the Levi-Civita connection associated to these local coordinates.
- (b) Show that the equator $\theta = \frac{\pi}{2}$ is the image of a geodesic. Are the parallels $\theta = \theta_0$ (with $\theta_0 \neq \frac{\pi}{2}$) images of geodesics?
- (c) Let $c : [0, 2\pi] \rightarrow S^2$ be the curve given in local coordinates by $(\theta(t), \varphi(t)) = (\theta_0, t)$, where $\theta_0 \in]0, \frac{\pi}{2}[$. Let V be a vector field parallel along c such that $V(0) = \frac{\partial}{\partial \theta}$. Compute the angle by which V is rotated when it returns to the point $c(2\pi) = c(0)$, that is, the angle between $V(0)$ and $V(2\pi)$.

(**Remark:** This is precisely the angle by which the oscillation plane of the Foucault pendulum – which is just any pendulum sufficiently long and heavy to remain oscillating for a long time – rotates after 24 hours at latitude $\frac{\pi}{2} - \theta_0$; the reason for this is that the oscillation plane tries to remain fixed with respect to the distant stars as the Earth rotates about its axis).

- (d) Indicate a geodesic triangle (that is, a triangle whose sides are images of geodesics) with 3 right angles. Without solving any differential equation, compute the angle by which a vector is rotated as it is parallel propagated once around the triangle.

2. Let (M, g) be a Riemannian manifold and $p \in M$. Show that the map $\mu : T_p M \rightarrow T_p^* M$ given by

$$\mu(v)(w) = g(v, w) \quad \forall v, w \in T_p M$$

is a linear isomorphism, given in the bases associated to local coordinates (x^1, \dots, x^n) by

$$\mu \left(\sum_{i=1}^n v^i \frac{\partial}{\partial x^i} \right) = \sum_{i,j=1}^n g_{ij} v^j dx^i.$$