

# Geometric Mechanics

2018/2019

1<sup>st</sup> Exam - 18 January 2019 - 13:00

1. Let  $M \subset \mathbb{R}^3$  be the surface of revolution given in cylindrical coordinates  $(\rho, \varphi, z)$  by

$$\rho = f(z),$$

where  $f : \mathbb{R} \rightarrow \mathbb{R}^+$  is a smooth function, and consider the metric induced on  $M$  by the Euclidean metric of  $\mathbb{R}^3$ ,

$$g_{\text{Euclidean}} = d\rho \otimes d\rho + \rho^2 d\varphi \otimes d\varphi + dz \otimes dz.$$

- (2/20) (a) Show that the equations for the geodesics of  $M$  are the Euler-Lagrange equations determined by the Lagrangian  $L : TM \rightarrow \mathbb{R}$  given in local coordinates by

$$L(\varphi, z, v^\varphi, v^z) = \frac{1}{2} ((f(z))^2 (v^\varphi)^2 + ((f'(z))^2 + 1) (v^z)^2).$$

- (2/20) (b) Check that the curves given in local coordinates by  $\varphi = \varphi_0$  (for any constant  $\varphi_0 \in \mathbb{R}$ ) or by  $f'(z) = 0$  are images of geodesics.

- (2/20) (c) Compute the Legendre transformation, show that  $L$  is hyper-regular and write an expression in local coordinates for the Hamiltonian  $H : T^*M \rightarrow \mathbb{R}$ .

- (2/20) (d) Show that  $H$  is completely integrable.

- (2/20) (e) Prove that if  $f'(z_0) = 0$  and  $f''(z_0) < 0$  then any geodesic whose image is  $z = z_0$  is stable, that is, if its initial condition is perturbed by a sufficiently small amount then the resulting geodesic remains close to  $z = z_0$ .

- (2/20) (f) Show that if  $f'(z_0) = 0$  and  $f''(z_0) > 0$  then there exist geodesics  $c : \mathbb{R} \rightarrow M$  accumulating on  $z = z_0$ , that is, satisfying  $z(c(t)) \neq z_0$  and  $\lim_{t \rightarrow +\infty} z(c(t)) = z_0$ .

- (2/20) (g) Consider now an ice skate moving on  $M$  in the particular case  $f(z) = 1$ , corresponding to the configuration space  $M \times S^1$ , kinetic energy

$$K(\varphi, z, \psi, v^\varphi, v^z, v^\psi) = \frac{m}{2} ((v^\varphi)^2 + (v^z)^2) + \frac{I}{2} (v^\psi)^2,$$

potential energy  $U = mgz$  (where  $m, I, g > 0$ ) and non-holonomic constraint

$$\Sigma = \text{span} \left\{ \cos \psi \frac{\partial}{\partial \varphi} + \sin \psi \frac{\partial}{\partial z}, \frac{\partial}{\partial \psi} \right\}.$$

Show that this is a true non-holonomic constraint, write the equations of motion and obtain a particular solution of these equations.

- (2/20) 2. Let  $M$  be an  $n$ -dimensional differentiable manifold,  $\omega \in \Omega^2(T^*M)$  the canonical symplectic form,  $H \in C^\infty(T^*M)$  a smooth function and  $E \in \mathbb{R}$  a regular value of  $H$ . Show that if  $G \in C^\infty(T^*M)$  satisfies  $\{G, H\} = 1$  then  $\Omega = \iota(X_G)\omega^n$  is a volume form on the  $(2n - 1)$ -dimensional submanifold  $H^{-1}(E) \subset T^*M$ , preserved by the flow of  $X_H$ , and independent of the choice of  $G$ .

(Remark: This volume form can be used to prove the Poicaré Recurrence Theorem on a single energy level; the existence of a local function  $G$  as above is guaranteed by Darboux's Theorem).

3. Recall that the Schwarzschild metric restricted to a radial line is given by

$$g_{\text{Schwarzschild}} = - \left(1 - \frac{2m}{r}\right) dt \otimes dt + \left(1 - \frac{2m}{r}\right)^{-1} dr \otimes dr,$$

where  $m > 0$  is the mass of the spherically symmetric object generating the field.

- (2/20) (a) Check that the equations for radial timelike geodesics can be written as

$$\begin{cases} \left(1 - \frac{2m}{r}\right) \frac{dt}{d\tau} = E \\ \left(\frac{dr}{d\tau}\right)^2 = E^2 - 1 + \frac{2m}{r} \end{cases},$$

where  $\tau$  is the proper time and  $E > 0$  is a constant.

- (2/20) (b) Show that a free-falling particle that is thrown upwards from  $r = r_0$ , reaches a maximum height  $r = r_1$ , and falls back to  $r = r_0$ , measures a proper time

$$\Delta\tau = 2 \int_{r_0}^{r_1} \frac{dr}{\sqrt{\frac{2m}{r} - \frac{2m}{r_1}}},$$

corresponding to a coordinate time

$$\Delta t = 2 \int_{r_0}^{r_1} \frac{\sqrt{1 - \frac{2m}{r_1}} dr}{\left(1 - \frac{2m}{r}\right) \sqrt{\frac{2m}{r} - \frac{2m}{r_1}}}.$$

Use this result to prove that

$$\Delta\tau > \Delta t \sqrt{1 - \frac{2m}{r_0}},$$

that is, the particle measures a longer proper time than the stationary observer sitting at  $r = r_0$ . (Hint: Compare the derivatives of  $\Delta\tau$  and  $\Delta t \sqrt{1 - \frac{2m}{r_0}}$  with respect to  $r_0$  using

$$\Delta t \geq 2 \int_{r_0}^{r_1} \frac{r_0 dr}{\sqrt{1 - \frac{2m}{r_1}} \sqrt{2m} \sqrt{r_1 - r}} \geq \frac{2r_0^2}{m \sqrt{1 - \frac{2m}{r_1}}} \sqrt{\frac{2m}{r_0} - \frac{2m}{r_1}}.$$