# Mathematical Relativity <br> 2021/2022 <br> $1^{\text {st }}$ Exam - 11 July 2022-10:00 

1. Let $(M, g)$ be a time-oriented spacetime and $p \in M$. For each of the following statements, give a proof if true or a counter-example if false.
(a) $I^{+}(p) \neq M$.
(b) $\partial I^{+}(p)=\partial J^{+}(p)$.
(c) If $M=\mathbb{R} \times S$ and $g=-d t^{2}+h$, where ( $S, h$ ) is a geodesically incomplete Riemannian manifold, then ( $M, g$ ) is not globally hyperbolic.
(d) If $q \in I^{+}(p)$ is connected to $p$ by a non-maximizing timelike geodesic $c$ then there exists a conjugate point along $c$ between $p$ and $q$.
2. The Nariai solution is the Lorentzian manifold ( $M, g$ ) given by $M=\mathbb{R}^{2} \times S^{2}$ and

$$
g=-d t^{2}+\cosh ^{2}(t) d x^{2}+d \theta^{2}+\sin ^{2} \theta d \varphi^{2}
$$

(where $(\theta, \varphi)$ are the usual spherical coordinates on $S^{2}$ ).
(a) Sketch the Penrose diagram of the Nariai solution.
(b) The Nariai solution solves the vacuum Einstein equations for a certain value of the cosmological constant $\Lambda$. Compute this value.
(c) Show that the Penrose singularity theorem does not hold if we replace "trapped surface" by "marginally trapped surface" in its statement (a marginally trapped surface is a surface whose null expansions are both nonpositive, as opposed to negative for a trapped surface).
3. The Kerr metric is given by

$$
\begin{aligned}
d s^{2}= & -\left(1-\frac{2 M r}{\rho^{2}}\right) d t^{2}-\frac{4 M a r \sin ^{2} \theta}{\rho^{2}} d t d \varphi+\frac{\rho^{2}}{\Delta} d r^{2} \\
& +\rho^{2} d \theta^{2}+\left(r^{2}+a^{2}+\frac{2 M a^{2} r \sin ^{2} \theta}{\rho^{2}}\right) \sin ^{2} \theta d \varphi^{2},
\end{aligned}
$$

where

$$
\begin{aligned}
& \rho^{2}=r^{2}+a^{2} \cos ^{2} \theta, \\
& \Delta=r^{2}-2 M r+a^{2}
\end{aligned}
$$

and $M, a \in \mathbb{R}$ are constants. Consider the region $r \gg M, a$.
(2/20) (a) Show that a positive orthonormal coframe is approximately given in this region by

$$
\begin{aligned}
& \omega^{0} \sim d t, \quad \omega^{r} \sim d r, \quad \omega^{\theta} \sim r d \theta \\
& \omega^{\varphi} \sim r \sin \theta d \varphi-\frac{2 M a \sin \theta}{r^{2}} d t
\end{aligned}
$$

(2/20)
(b) Setting $X=\frac{\partial}{\partial t}$ and $Y=\frac{\partial}{\partial \varphi}$, establish the following asymptotic formulas:

$$
\begin{aligned}
& X^{b} \sim-\left(1-\frac{2 M}{r}\right) d t-\frac{2 M a \sin ^{2} \theta}{r} d \varphi ; \\
& Y^{b} \sim-\frac{2 M a \sin ^{2} \theta}{r} d t+r^{2} \sin ^{2} \theta d \varphi ; \\
& d X^{b} \sim \frac{2 M}{r^{2}} \omega^{0} \wedge \omega^{r}+\ldots ; \\
& d Y^{b} \sim-\frac{6 M a \sin ^{2} \theta}{r^{2}} \omega^{0} \wedge \omega^{r}+\ldots
\end{aligned}
$$

(2/20) (c) Prove that the Komar mass and the Komar angular momentum of the Kerr solution are given by $M_{\text {Komar }}=M$ and $J_{\text {Komar }}=M a$.

