

# Mathematical Relativity

## Homework 2

*Due on March 2*

Consider the spherically symmetric Lorentzian metric given by

$$g = -(A(t, r))^2 dt^2 + (B(t, r))^2 dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\varphi^2,$$

where  $A$  and  $B$  are positive smooth functions.

1. Use the condition of compatibility with the metric and Cartan's first structure equations,

$$\begin{cases} \omega_{\mu\nu} = -\omega_{\nu\mu} \\ d\omega^\mu + \omega^\mu{}_\nu \wedge \omega^\nu = 0 \end{cases}$$

to show that the nonvanishing connection forms for the orthonormal frame dual to

$$\omega^0 = A dt, \quad \omega^r = B dr, \quad \omega^\theta = r d\theta, \quad \omega^\varphi = r \sin \theta d\varphi$$

are (using the notation  $\dot{\phantom{x}} = \frac{\partial}{\partial t}$  and  $\phantom{x}' = \frac{\partial}{\partial r}$ )

$$\begin{aligned} \omega^0{}_r = \omega^r{}_0 &= \frac{A'}{B} dt + \frac{\dot{B}}{A} dr; \\ \omega^\theta{}_r = -\omega^r{}_theta &= \frac{1}{B} d\theta; \\ \omega^\varphi{}_r = -\omega^r{}_phi &= \frac{\sin \theta}{B} d\varphi; \\ \omega^\varphi{}_theta = -\omega^\theta{}_phi &= \cos \theta d\varphi. \end{aligned}$$

2. Use Cartan's second structure equations

$$\Omega^\mu{}_\nu = d\omega^\mu{}_\nu + \omega^\mu{}_\alpha \wedge \omega^\alpha{}_\nu$$

to show that the curvature forms on this frame are

$$\begin{aligned} \Omega^0{}_r = \Omega^r{}_0 &= \left( \frac{A''B - A'B'}{AB^3} + \frac{\dot{A}\dot{B} - A\ddot{B}}{A^3B} \right) \omega^r \wedge \omega^0; \\ \Omega^0{}_theta = \Omega^\theta{}_0 &= \frac{A'}{rAB^2} \omega^\theta \wedge \omega^0 + \frac{\dot{B}}{rAB^2} \omega^\theta \wedge \omega^r; \\ \Omega^0{}_phi = \Omega^\phi{}_0 &= \frac{A'}{rAB^2} \omega^\varphi \wedge \omega^0 + \frac{\dot{B}}{rAB^2} \omega^\varphi \wedge \omega^r; \\ \Omega^\theta{}_r = -\Omega^r{}_theta &= \frac{B'}{rB^3} \omega^\theta \wedge \omega^r + \frac{\dot{B}}{rAB^2} \omega^\theta \wedge \omega^0; \\ \Omega^\varphi{}_r = -\Omega^r{}_phi &= \frac{B'}{rB^3} \omega^\varphi \wedge \omega^r + \frac{\dot{B}}{rAB^2} \omega^\varphi \wedge \omega^0; \\ \Omega^\varphi{}_theta = -\Omega^\theta{}_phi &= \frac{B^2 - 1}{r^2 B^2} \omega^\varphi \wedge \omega^\theta. \end{aligned}$$

3. Using

$$\Omega^\mu{}_\nu = \sum_{\alpha < \beta} R_{\alpha\beta}{}^\mu{}_\nu \omega^\alpha \wedge \omega^\beta$$

determine the components  $R_{\alpha\beta}{}^\mu{}_\nu$  of the curvature tensor on this orthonormal frame, and show that the nonvanishing components of the Ricci tensor on this frame are

$$\begin{aligned} R_{00} &= \frac{A''B - A'B'}{AB^3} + \frac{\dot{A}\dot{B} - A\ddot{B}}{A^3B} + \frac{2A'}{rAB^2}; \\ R_{0r} &= R_{r0} = \frac{2\dot{B}}{rAB^2}; \\ R_{rr} &= \frac{A'B' - A''B}{AB^3} + \frac{A\ddot{B} - \dot{A}\dot{B}}{A^3B} + \frac{2B'}{rB^3}; \\ R_{\theta\theta} &= R_{\varphi\varphi} = -\frac{A'}{rAB^2} + \frac{B'}{rB^3} + \frac{B^2 - 1}{r^2B^2}. \end{aligned}$$

Conclude that the nonvanishing components of the Einstein tensor on this frame are

$$\begin{aligned} G_{00} &= \frac{2B'}{rB^3} + \frac{B^2 - 1}{r^2B^2}; \\ G_{0r} &= G_{r0} = \frac{2\dot{B}}{rAB^2}; \\ G_{rr} &= \frac{2A'}{rAB^2} - \frac{B^2 - 1}{r^2B^2}; \\ G_{\theta\theta} &= G_{\varphi\varphi} = \frac{A''B - A'B'}{AB^3} + \frac{\dot{A}\dot{B} - A\ddot{B}}{A^3B} + \frac{A'}{rAB^2} - \frac{B'}{rB^3}. \end{aligned}$$

4. Show that if we write

$$B(t, r) = \left(1 - \frac{2m(t, r)}{r}\right)^{-\frac{1}{2}}$$

for some smooth function  $m$  then

$$G_{00} = \frac{2m'}{r^2}.$$

Conclude that the vacuum Einstein equations  $G_{00} = G_{0r} = 0$  are equivalent to

$$B = \left(1 - \frac{2M}{r}\right)^{-\frac{1}{2}},$$

where  $M \in \mathbb{R}$  is an integration constant.

5. Show that the vacuum equation  $G_{00} + G_{rr} = 0$  is equivalent to  $A = \frac{\alpha(t)}{B}$  for some positive smooth function  $\alpha(t)$ .
6. Check that if  $A$  and  $B$  are as above then the remaining vacuum equation  $G_{\theta\theta} = G_{\varphi\varphi} = 0$  is automatically satisfied.
7. Argue that it is always possible to rescale the coordinate  $t$  so that the any metric of the given form satisfying the vacuum Einstein field equations is written

$$g = -\left(1 - \frac{2M}{r}\right) dt^2 + \left(1 - \frac{2M}{r}\right)^{-1} dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\varphi^2.$$