Mathematical Relativity

Homework 13

Due on May 25

1. (a) Check that the Klein-Gordon equation in Minkowski spacetime, $\Box \phi - m^2 \phi = 0$, can be derived from the Lagrangian density

$$\mathcal{L} = -\frac{1}{2} \left(\partial_{\mu} \phi \, \partial^{\mu} \phi + m^2 \phi^2 \right).$$

(b) Show that the corresponding Hamiltonian is

$$H = \int_{\mathbb{R}^n} \frac{1}{2} \left((\partial_0 \phi)^2 + \ldots + (\partial_n \phi)^2 + m^2 \phi^2 \right) dx^1 \wedge \ldots \wedge dx^n.$$

(c) Starting with the Einstein-Hilbert-Klein-Gordon action

$$S = \int_{M} \left[R - 8\pi \left(g^{\mu\nu} \partial_{\mu} \phi \, \partial_{\nu} \phi + m^{2} \phi^{2} \right) \right] dV_{n+1}$$

obtain the energy-momentum tensor for ϕ :

$$T_{\mu\nu} = \partial_{\mu}\phi \,\partial_{\nu}\phi - \frac{1}{2}g_{\mu\nu} \left(\partial_{\alpha}\phi \,\partial^{\alpha}\phi + m^2\phi^2\right).$$

- (d) Check that T_{00} coincides with the Hamiltonian density \mathcal{H} (i.e. the integrand in the expression for the Hamiltonian H).
- 2. Let γ be the spherically symmetric Riemannian metric defined in \mathbb{R}^3 by

$$\gamma = \frac{dr^2}{1 - \frac{2m(r)}{r}} + r^2 \left(d\theta^2 + \sin^2 \theta d\varphi^2 \right),$$

where m is a smooth function whose derivative has compact support.

(a) Check that in Cartesian coordinates we have

$$\gamma_{ij} = \delta_{ij} + \frac{\frac{2m(r)}{r^3}}{1 - \frac{2m(r)}{r}} x^i x^j.$$

(b) Show that if the limit

$$M = \lim_{r \to \infty} m(r)$$

exists then γ is asymptotically flat with ADM mass M (which in particular coincides with the Komar mass when appropriate).

(c) Check that γ has scalar curvature

$$R = \frac{4}{r^2} \frac{dm}{dr},$$

and use this to prove the Riemannian Positive Mass Theorem for $\gamma. \\$

(d) Show that $r=r_0$ is a minimal surface if and only if $m(r_0)=\frac{r_0}{2}$ (in which case r is a well-defined coordinate only for $r>r_0$), and use this to prove the Riemannian Penrose Conjecture for γ .